

## RECOMMENDATION ITU-R P.1147-1

**PREDICTION OF SKY-WAVE FIELD STRENGTH AT FREQUENCIES  
BETWEEN ABOUT 150 AND 1700 kHz**

(Question ITU-R 225/3)

(1995-1999)

The ITU Radiocommunication Assembly,

*considering*

- a) that there is a need to give guidance to engineers in the planning of broadcast services in the LF and MF bands;
- b) that it is important, for stations working in the same or adjacent frequency channels, to determine the minimum geographical separation required to avoid interference resulting from long-distance ionospheric propagation;
- c) that portions of this frequency range are being shared by broadcasting and other services in different Regions, an accurate method for predicting interference levels is needed to maintain efficient and orderly utilization of those portions of the spectrum,

*recommends*

that the following method be adopted for use, taking particular note of the discussion on accuracy given in Annex 1.

## 1 Introduction

This method predicts values of the night-time sky-wave field strength for a given power radiated from one or more vertical antennas, when measured by a loop antenna at ground level aligned in a vertical plane along the great-circle path to the transmitter. The method has been based on measurements made in the frequency bands allocated to broadcasting and applies for paths of length 50 to 12 000 km for those LF and MF bands in particular. For a discussion on daytime propagation, see Annex 2.

Figures 1, 2 and 3 are an essential part of the prediction method. Geomagnetic maps are included for convenience in Figs. 5, 6 and 10. The remaining Figures and Appendix 1 provide additional information to simplify the use of the method.

## 2 Annual median night-time field strength

The predicted sky-wave field strength is given by:

$$E = V + E_0 - L_t = V + G_S - L_p + A - 20 \log p - L_a - L_t - L_r \quad (1)$$

where:

$E$ : annual median of half-hourly median field strengths (dB( $\mu$ V/m)) for a given transmitter cymomotive force,  $V$ , and at a given time,  $t$ , relative to sunset or sunrise as appropriate

$E_0$ : annual median of half-hourly median field strengths (dB( $\mu$ V/m)) for a transmitter cymomotive force of 300 V at the reference time defined in § 2.1

$V$ : transmitter cymomotive force (dB above a reference cymomotive force of 300 V) (see § 2.2)

$G_S$ : sea-gain correction (dB) (see § 2.3)

$L_p$ : excess polarization-coupling loss (dB) (see § 2.4)

$A$ : a constant. At LF,  $A = 110.2$ . At MF,  $A = 107$  except for propagation paths whose midpoints are situated in the part of Region 3 south of parallel  $11^\circ$  S. In those cases,  $A = 110$

$L_a$ : loss factor incorporating effects of ionospheric absorption and related factors (see § 2.6)

$L_t$ : hourly loss factor (dB) (see § 2.7)

$L_r$ : loss factor incorporating effect of solar activity (§ 2.8).

Figure 4 shows  $E_0$  as a function of ground distance,  $d$ , for various geomagnetic latitudes when  $G_S$ ,  $L_p$  and  $R$  are all zero; where  $R$  is the twelve-month smoothed international relative sunspot number.

## 2.1 Reference time

The reference time is taken as six hours after the time at which the Sun sets at a point S on the surface of the Earth. For paths shorter than 2 000 km, S is the mid-point of the path. On longer paths, S is 750 km from the terminal where the Sun sets last, measured along the great-circle path.

## 2.2 Cymomotive force

The transmitter cymomotive force  $V$  (dB(300 V)) is given as:

$$V = P + G_V + G_H \quad (2)$$

where:

$P$ : radiated power (dB(1 kW))

$G_V$ : transmitting antenna gain factor (dB) due to vertical directivity, given in Fig. 1

$G_H$ : transmitting antenna gain factor (dB) due to horizontal directivity. For directional antennas,  $G_H$  is a function of azimuth. For omnidirectional antennas,  $G_H = 0$ .

## 2.3 Sea gain

The sea gain  $G_S$  is the additional signal gain when one or both terminals is situated near the sea, but it does not apply to propagation over fresh water.  $G_S$  for a single terminal is given by:

$$G_S = G_0 - c_1 - c_2 \quad \text{for} \quad (c_1 + c_2) < G_0 \quad (3)$$

$$G_S = 0 \quad \text{for} \quad (c_1 + c_2) \geq G_0 \quad (4)$$

where:

$G_0$ : sea gain when the terminal is on the coast and the path is unobstructed by further land (dB)

$c_1$ : correction to take account of the distance between the terminal and the sea

$c_2$ : correction to take account of the width of one or more sea channels, or the presence of islands.

If both terminals are near the sea,  $G_S$  is the sum of the values for the individual terminals.

$G_0$  is given in Fig. 2 as a function of  $d$  for LF and MF. At MF,  $G_0 = 10$  dB when  $d > 6 500$  km; and at LF  $G_0 = 4.1$  dB when  $d > 5 000$  km, where  $d$  is the ground distance between the two terminals.

The correction  $c_1$  is given by:

$$c_1 = \frac{s_1}{r_1} G_0 \quad (5)$$

where:

$s_1$ : distance of terminal from sea, measured along great-circle path (km)

$$r_1 = 10^3 G_0^2 / Q_1 f \quad \text{km}$$

$f$ : frequency (kHz)

$Q_1 = 0.30$  at LF and 1.4 at MF.

The correction  $c_2$  is given by:

$$c_2 = \alpha G_0 \left( 1 - \frac{s_2}{r_2} \right) \quad \text{for} \quad s_2 < r_2 \quad (6)$$

$$c_2 = 0 \quad \text{for} \quad s_2 \geq r_2 \quad (7)$$

where:

$s_2$ : distance of terminal from next section of land, measured along the great-circle path (km)

$$r_2 = 10^3 G_0^2 / Q_2 f \quad \text{km}$$

$Q_2 = 0.25$  at LF and 1.2 at MF

$\alpha$ : proportion of land in the section of path between  $r_2$  and  $s_2$  ( $0 < \alpha \leq 1$ ).

If a computer is used but a terrain data bank is not available to calculate  $\alpha$ , then  $\alpha$  should be made equal to 0.5, which implies that land and sea are present in equal proportions in the section of path between  $r_2$  and  $s_2$ .

To facilitate calculation, Fig. 8a shows  $r_1$ , the greatest distance from the sea for which sea gain has to be calculated, and Fig. 8b shows  $r_2$ , the greatest distance to the next section of land for which the correction  $c_2$  is required, for various frequencies.

## 2.4 Polarization coupling loss

$L_p$  is the excess polarization coupling loss (dB). At LF,  $L_p = 0$ . At MF,  $L_p$  for a single terminal is given by one of the following two formulae:

$$\begin{aligned} L_p &= 180 (36 + \theta^2 + I^2)^{-1/2} - 2 \quad \text{dB} \quad \text{for} \quad I \leq 45^\circ \\ L_p &= 0 \quad \text{for} \quad I > 45^\circ \end{aligned} \quad (8)$$

where  $I$  is the magnetic dip, N or S (degrees) at the terminal and  $\theta$  is the path azimuth measured in degrees from the magnetic E-W direction, such that  $|\theta| \leq 90^\circ$ .  $L_p$  should be evaluated separately for the two terminals, because of the different values of  $\theta$  and  $I$  that may apply, and the two  $L_p$  values added. The most accurate available values of magnetic dip and declination (e.g. see Figs. 5 and 6) should be used in determining  $\theta$  and  $I$ .

Figure 7 shows values of  $L_p$  calculated from equation (8).

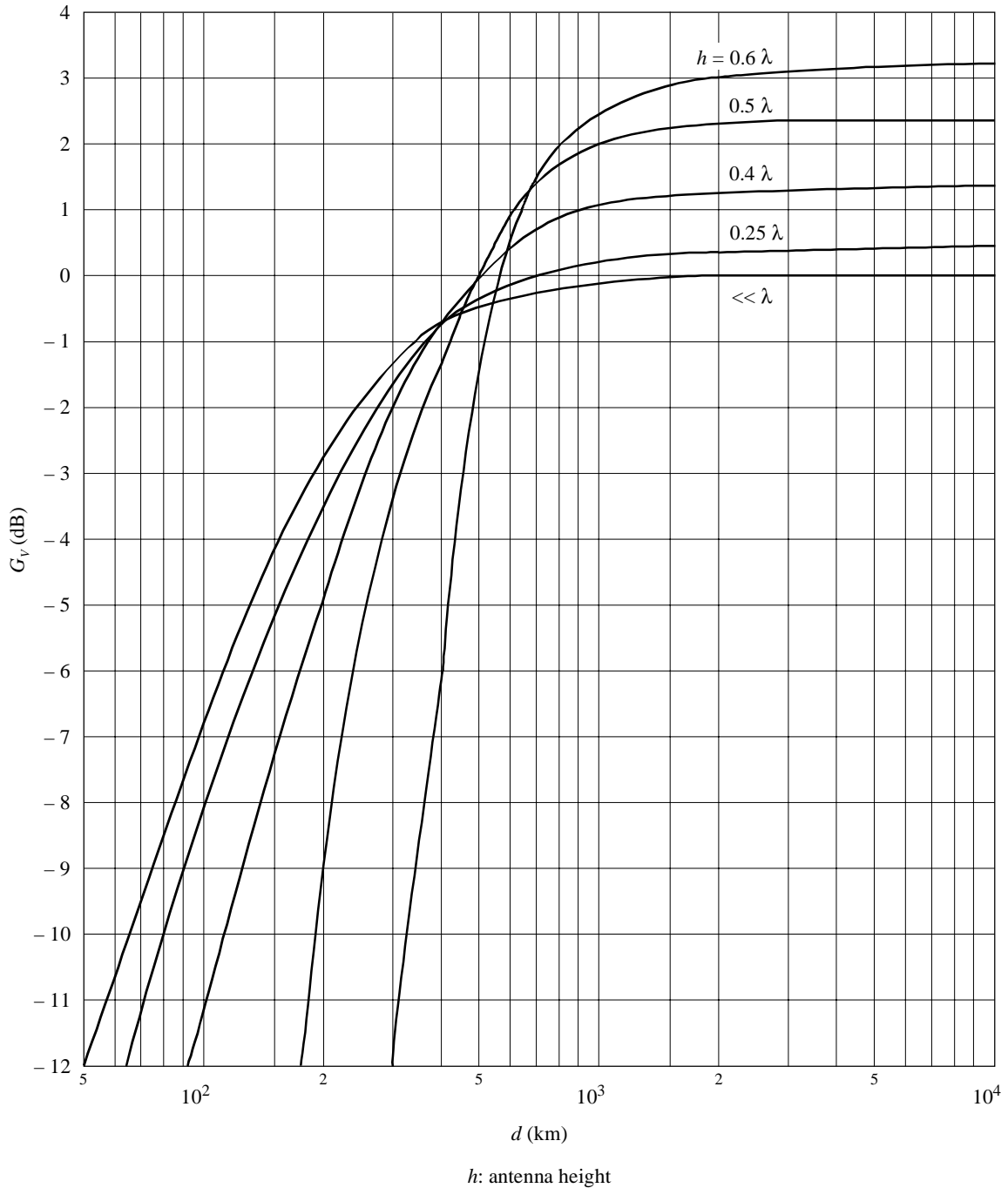
## 2.5 Slant propagation distance, $p$

For paths longer than 1 000 km,  $p$  (km) is approximately equal to the ground distance,  $d$  (km) between transmitter and receiver. For shorter paths:

$$p = (d^2 + 40\,000)^{1/2} \quad (9)$$

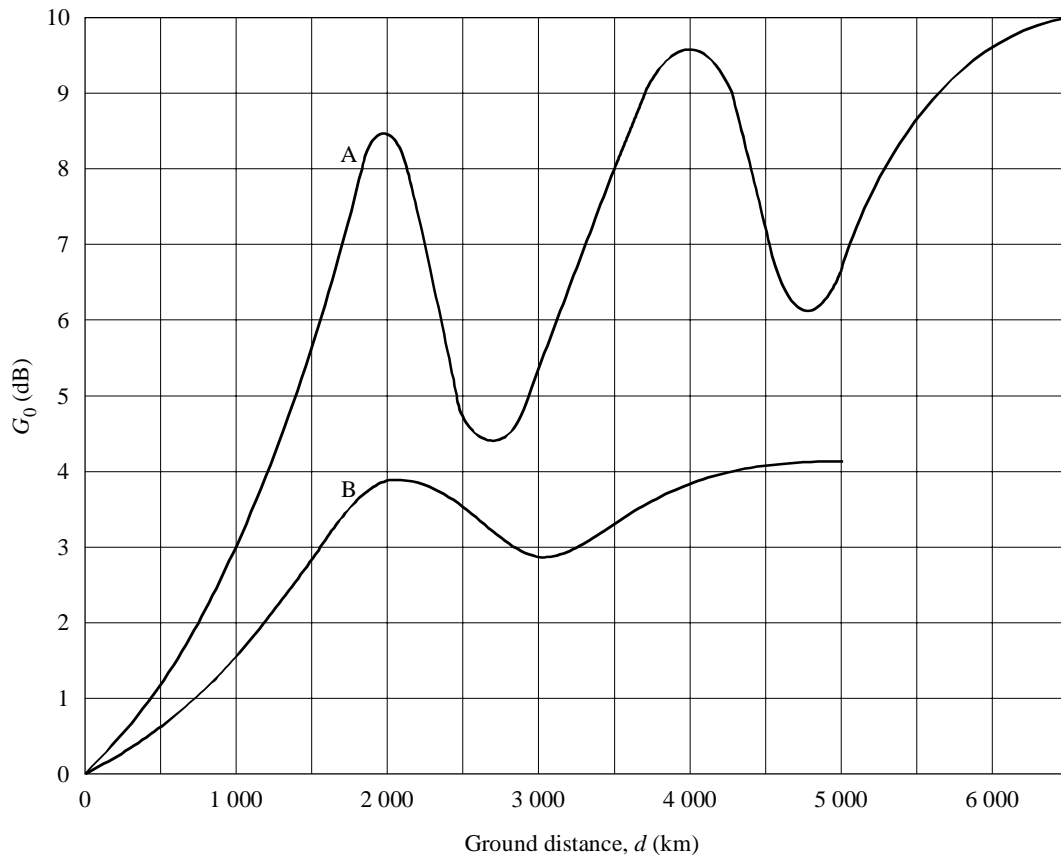
Equation (9) may be used for paths of any length with negligible error. It should be used in all cases where the distances considered are both above and below 1 000 km, to avoid discontinuities in field strength as a function of distance.

FIGURE 1  
 Transmitting antenna gain factor for single monopoles ( $G_v$ )  
 over perfect Earth



Note 1 – For  $10\ 000\ \text{km} < d < 12\ 000\ \text{km}$ , readings at  $d = 10\ 000\ \text{km}$  are to be used.

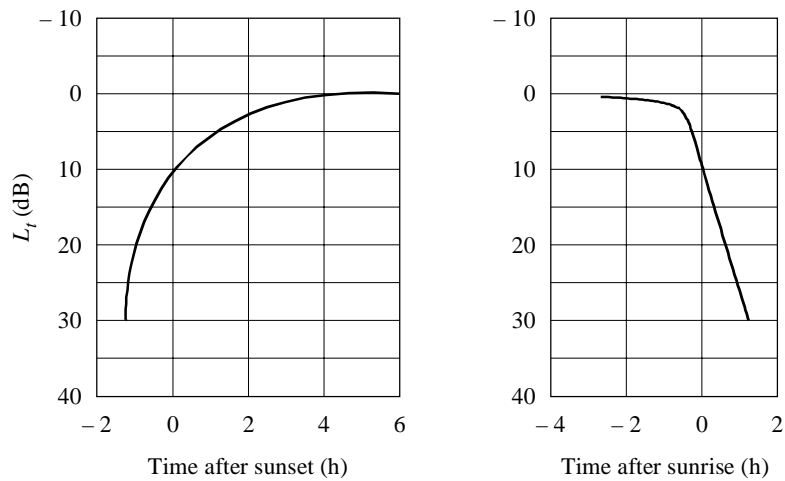
FIGURE 2  
Sea gain ( $G_0$ ) for a single terminal on the coast



Curves A: MF band  
B: LF band

1147-02

FIGURE 3  
Hourly loss factor ( $L_t$ )



1147-03

## 2.6 Loss factor incorporating effects of ionospheric absorption and related factors

$L_a$  is the loss factor incorporating effects of ionospheric absorption, focusing, terminal losses and loss between hops of multi-hop paths.

$$L_a = k \sqrt{p / 1000} \quad (10)$$

and the basic loss coefficient  $k$  is given by:

$$k = (2\pi + 4.95 \tan^2 \Phi) \quad (11)$$

In equation (11)  $\Phi$  is the (dipole) geomagnetic latitude of the mid-point of the path under study. Northern latitudes are considered positive, southern latitudes negative. Figure 10 may be used to obtain the geomagnetic latitude from known geographical coordinates of the mid-point. If the value of  $\Phi$  is greater than  $+60^\circ$ , equation (11) is evaluated for  $\Phi = +60^\circ$ . If the value of  $\Phi$  is less than  $-60^\circ$ , equation (11) is evaluated for  $\Phi = -60^\circ$ . Paths longer than 3 000 km are divided into two equal sections,  $k$  for each section is determined according to equation (11). The algebraic average of the two values of  $k$  ( $k_{eff}$ ) is then used in calculating field strength (see also Fig. 9).

## 2.7 Hourly loss factor

The hourly loss factor,  $L_t$  (dB) is given in Fig. 3. The time  $t$  is the time in hours relative to the sunrise or sunset time as appropriate. These are taken at the ground at the mid-path position for  $d < 2000$  km and at 750 km from the terminal where the Sun sets or rises first for longer paths. The large values of hourly loss factor near midday are not defined (Fig. 3). For times in this period use a limit value of 30 dB. The hourly loss factor should not be calculated for high latitude paths and seasons when sunrise and sunset do not occur.

Equations generally equivalent to these curves to within 0.5 dB are given in § 1 of Appendix 1. Figure 3 represents the average annual diurnal variation.

Figure 11 shows sunset and sunrise times for a range of geographic latitudes and months. Equations equivalent to these sunset and sunrise curves are given in § 2 of Appendix 1.

## 2.8 Loss factor incorporating effect of solar activity

$L_r$  is the loss factor incorporating effect of solar activity and is given by the following formulae:

at LF,		$L_r = 0$	
at MF,	if $ \Phi  \leq 45^\circ$ ,	$L_r = 0$	
at MF,	if $ \Phi  > 45^\circ$ ,	$L_r = b(R/100) (p/1000)$	dB

where:

$$b = (|\Phi| - 45)/3 \quad \text{except in Europe where } b = 1 \text{ is to be used regardless of latitude.} \quad (13)$$

Paths longer than 3 000 km are divided into two equal sections as described in § 2.6. The value of  $L_r$  for each section is derived and added together.

### 3 Day-to-day and short-period variations of night-time field strengths

The difference,  $\Delta$ , at a specific time relative to sunset or sunrise, between the field strength exceeded for 10% of the time and the annual median value is given by:

at LF,  $\Delta = 6.5$  dB (14a)

at MF,  $\Delta = 0.2 |\Phi| - 2$  dB (14b)

$\Delta$  is greater than or equal to 6 dB but less than or equal to 10 dB.

FIGURE 4a  
 Curves showing  $E_0$  for LF when  $G_s, L_p$  and  $R$  are all zero, for constant geomagnetic latitudes

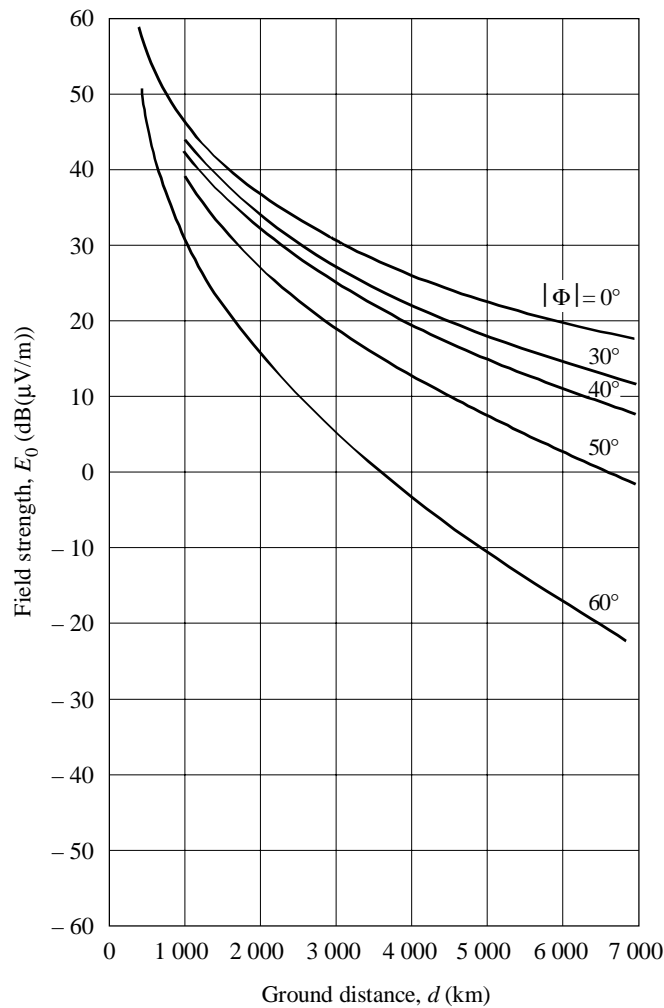
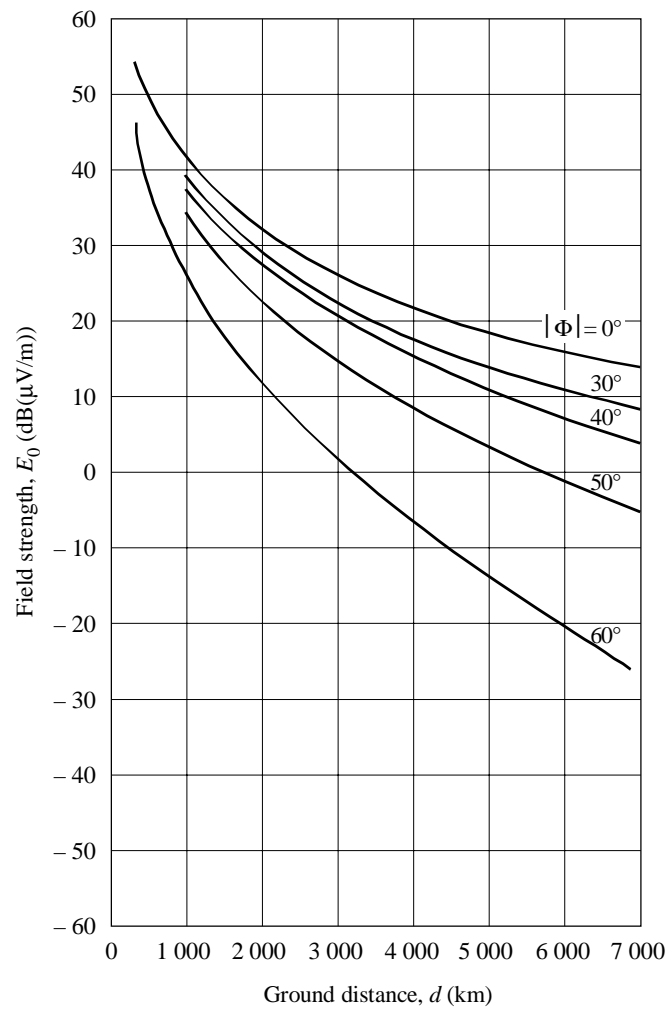


FIGURE 4b  
 Curves showing  $E_0$  for MF when  $G_s$ ,  $L_p$  and  $R$  are  
 all zero, for constant geomagnetic latitudes

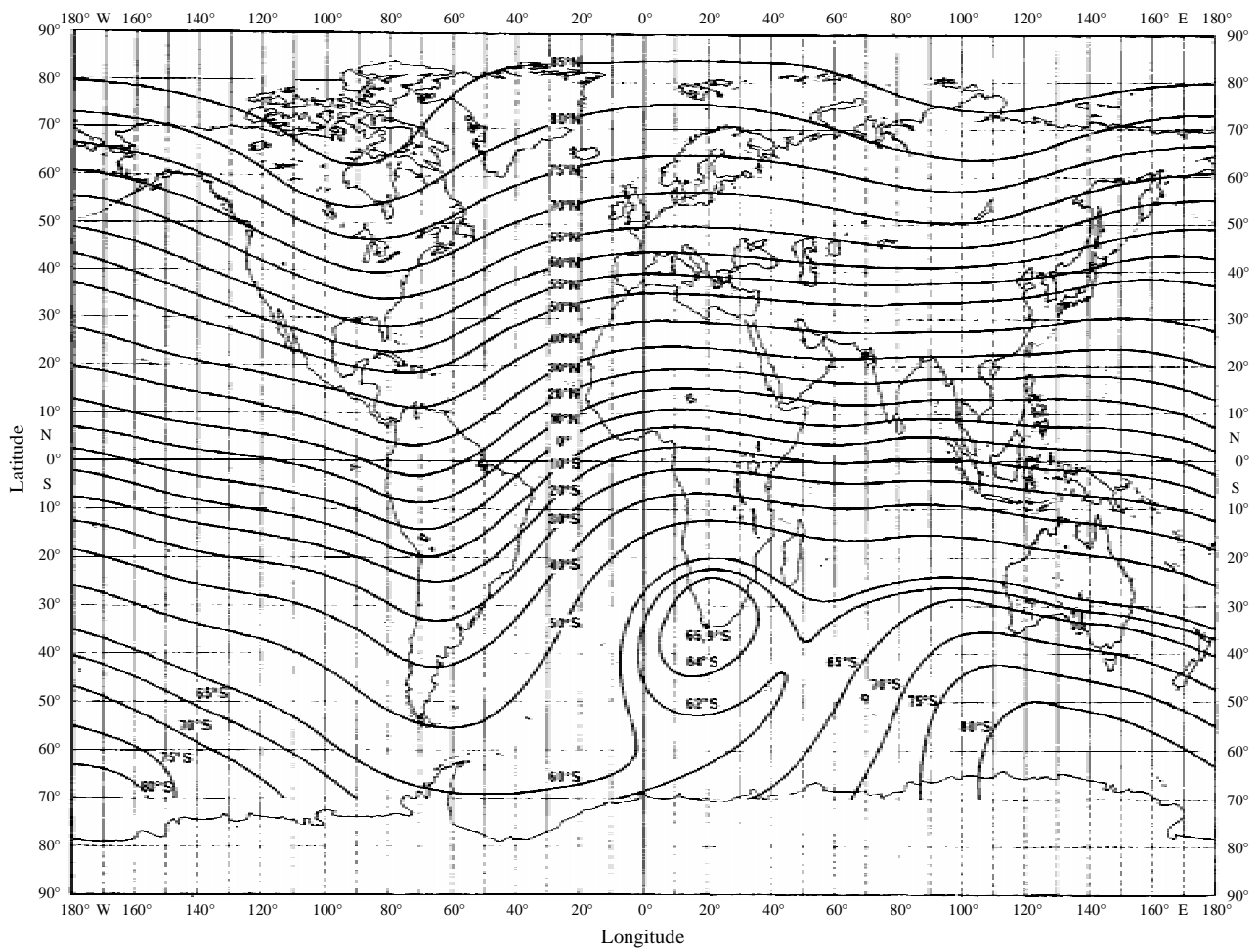


Note 1 – For applications in Southern Region 3, add 3 dB.

1147-04b



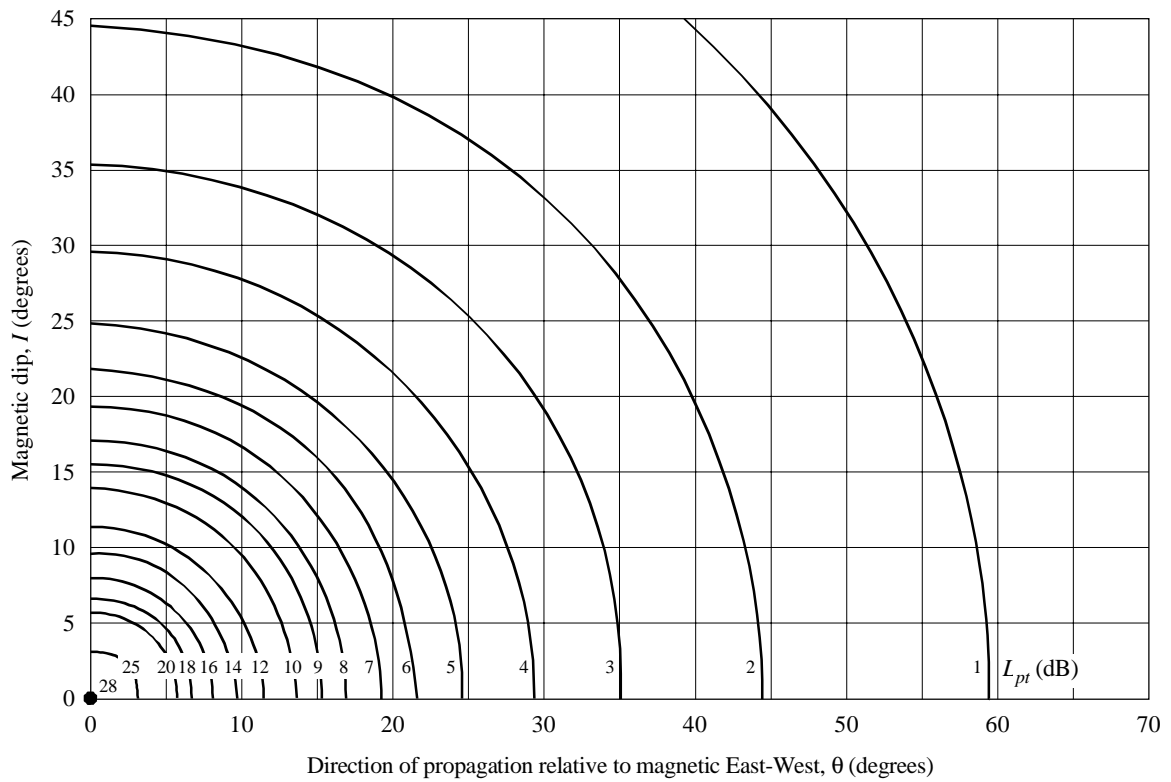
FIGURE 5  
Map of magnetic dip (epoch 1975.0)



(Source: Magnetic inclination or dip (epoch 1975.0) Chart No. 30 World U.S. Defense Mapping Agency Hydrographic Center)



FIGURE 7  
 Excess polarization coupling loss  $L_p$  (for a single terminal)



$$L_p = 180 (36 + \theta^2 + I^2)^{-1/2} - 2$$

1147-07

FIGURE 8a  
Values of  $r_1$  for various frequencies

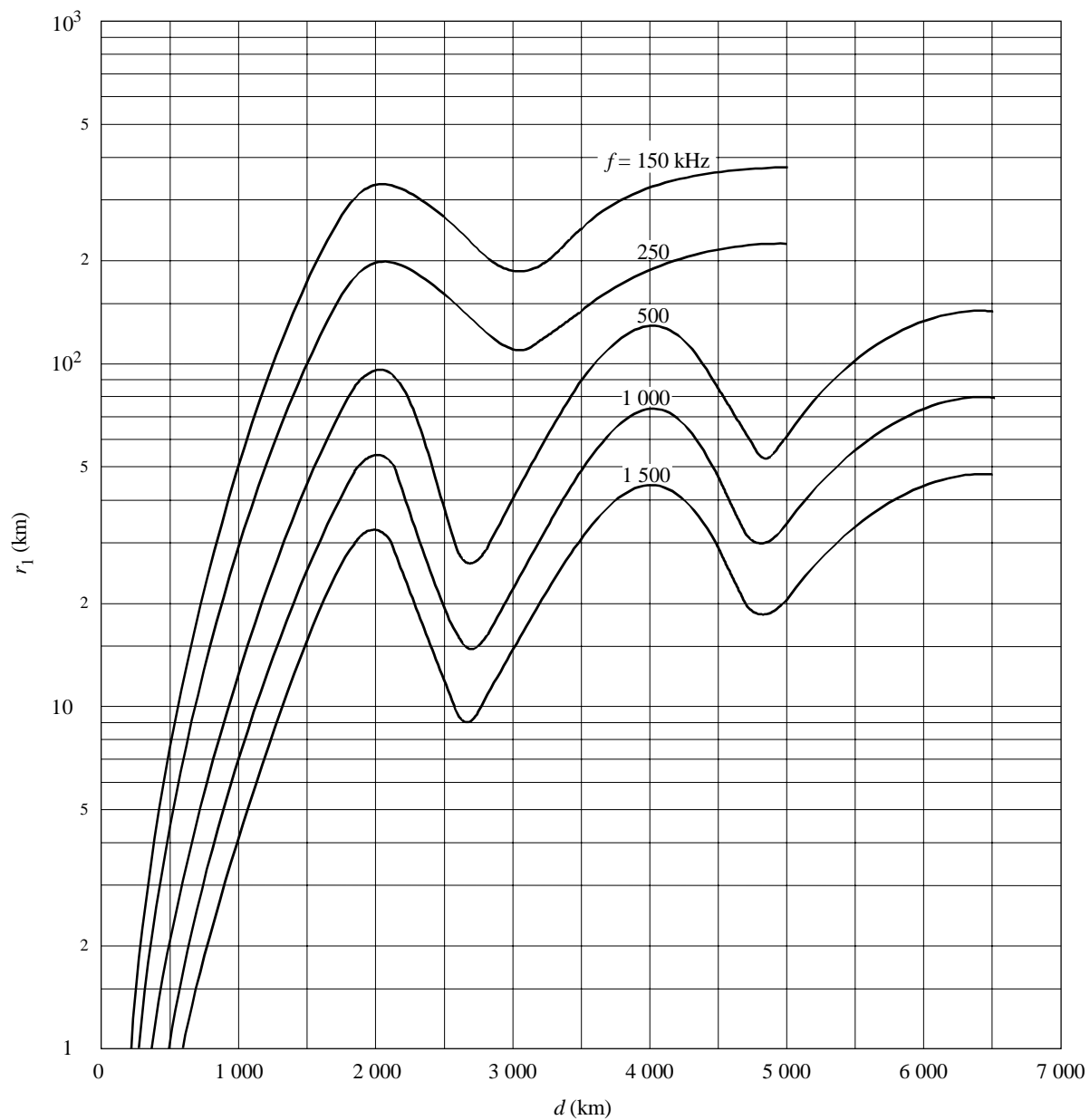


FIGURE 8b  
 Values of  $r_2$  for various frequencies

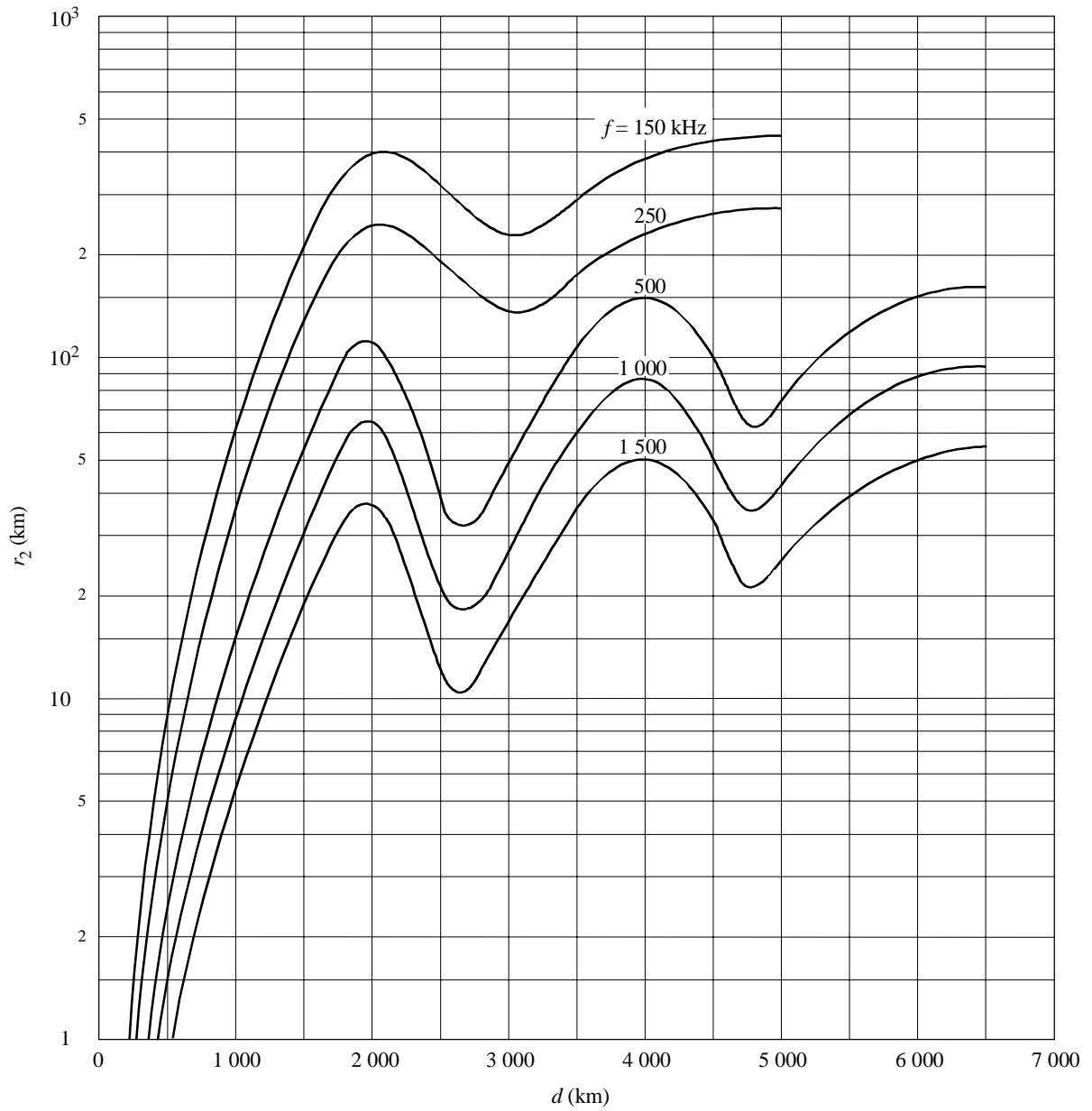
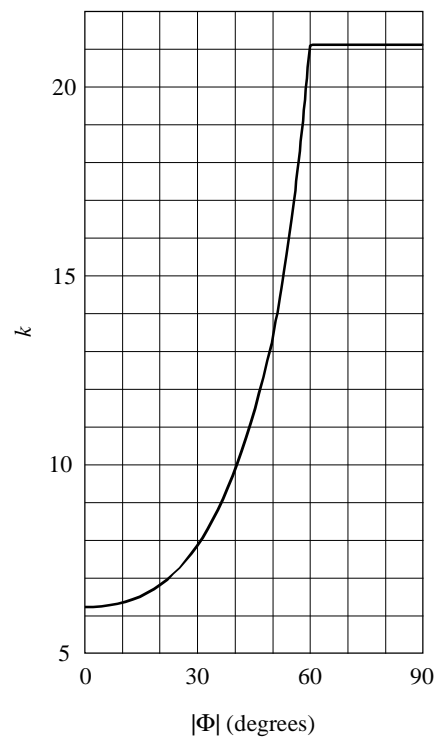


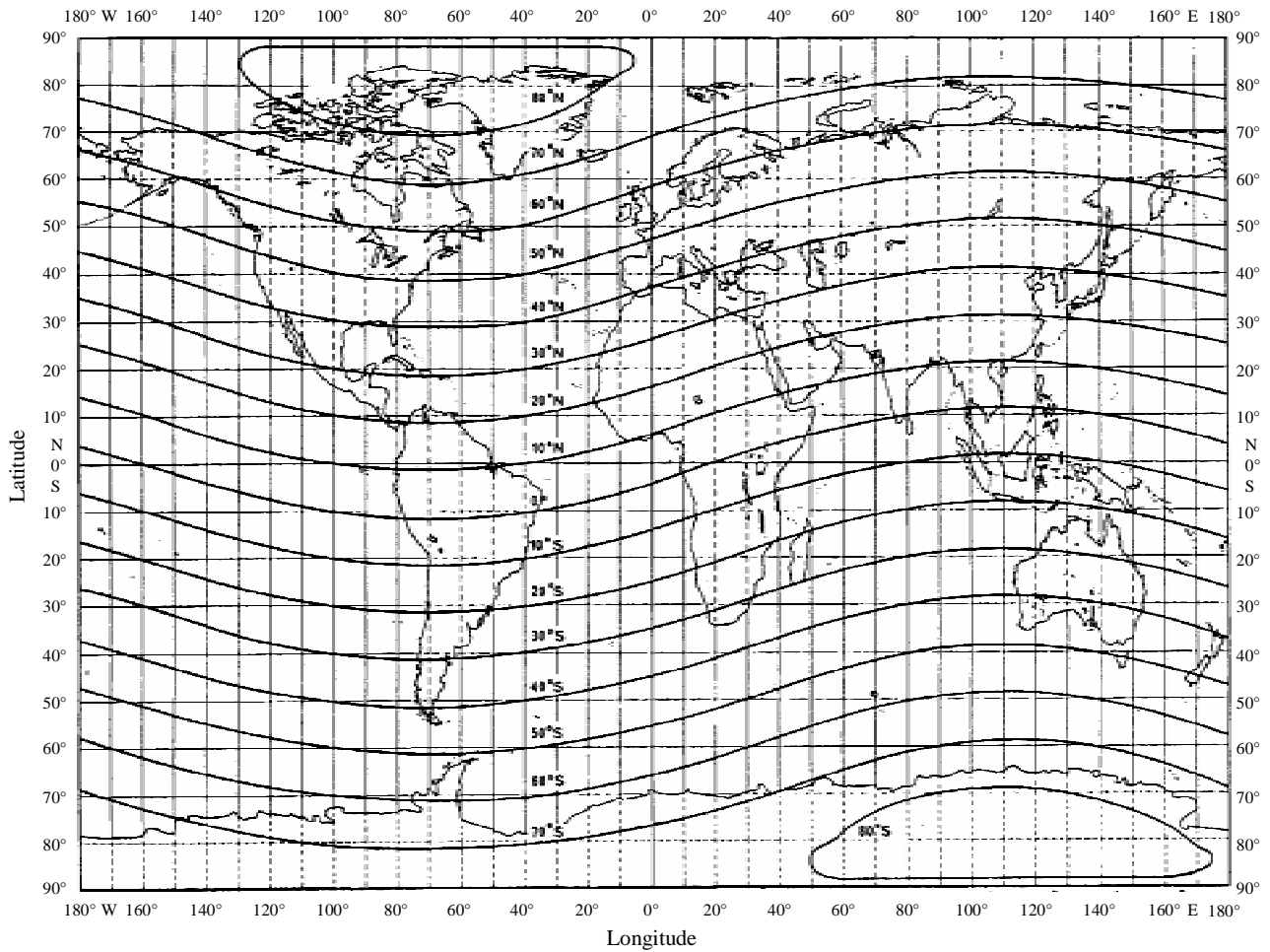
FIGURE 9  
Basic loss coefficient



$$k = (2\pi + 4.95 \tan^2 \Phi) \quad \text{for } |\Phi| \leq 60^\circ$$

1147-09

FIGURE 10  
Geomagnetic latitudes



$$\Phi = \arcsin [\sin \alpha \cdot \sin 78.5^\circ + \cos \alpha \cdot \cos 78.5^\circ \cdot \cos (69^\circ + \beta)]$$

$\Phi$ : geomagnetic latitude

$\alpha$ : geographic latitude

$\beta$ : geographic longitude

North and East coordinates are considered positive, and South and West coordinates negative.





## 1 Hourly loss factor: $L_t$

These equations may be used instead of the curves in Fig. 3, within the stated limits of  $t$ . For hours between these times (i.e. near midnight) set  $L_t = 0$ .

$$\begin{aligned} L_t \text{ (sunset)} &= 12.40 - 9.248 t + 2.892 t^2 - 0.3343 t^3 && \text{for } -1 < t \text{ (sunset)} < 4 \\ L_t \text{ (sunrise)} &= 9.6 + 12.2 t + 5.62 t^2 + 0.86 t^3 && \text{for } -3 < t \text{ (sunrise)} < 1 \end{aligned}$$

where  $t$  is the time in hours relative to sunset or sunrise at the path mid-point.

## 2 Sunset and sunrise times

For non-polar locations, i.e. such that  $|\alpha| < 65^\circ$ , the times of sunset and sunrise can be calculated as follows, to an accuracy of  $\pm 2$  min:

$N$ : day of year, in days; e.g. 1 January = 1

$S'$ : approximate local time of event, e.g. sunset = 1800 h, sunrise = 0600 h

$Z$ : Sun's zenith distance (degrees) =  $90.8333^\circ(90^\circ 50')$  for sunset or sunrise.

*Step 1:* Calculate observer's longitude,  $B$ :

$$B = \beta / 15 \quad \text{h}$$

*Step 2:* Calculate the time of event,  $Y$ :

$$Y = N + (S' - B) / 24 \quad \text{days}$$

*Step 3:* Calculate Sun's mean anomaly,  $M$ :

$$M = 0.985600 Y - 3.289 \quad \text{degrees}$$

*Step 4:* Calculate Sun's longitude,  $L$ :

$$L = M + 1.916 \sin M + 0.020 \sin 2 M + 282.634 \quad \text{degrees}$$

Note in which quadrant  $L$  occurs.

*Step 5:* Calculate Sun's right ascension,  $RA$ :

$$\tan RA = 0.91746 \tan L$$

Note that  $RA$  must be in the same quadrant as  $L$ .

*Step 6:* Calculate Sun's declination,  $s$ :

$\sin s = 0.39782 \sin L$ , from which:

$$\cos s = + \sqrt{1 - \sin^2 s}$$

Note that  $\sin s$  may be positive or negative but  $\cos s$  must always be positive.

*Step 7:* Calculate Sun's local hour angle,  $H$ :

$$\cos H = x = (\cos Z - \sin s \cdot \sin \alpha) / (\cos s \cdot \cos \alpha)$$

Note that if  $|x| > 1$ , there is no sunset or sunrise.

From  $\cos H$ , obtain  $H$  in degrees; for sunrise  $180 < H < 360$ ; for sunset  $0 < H < 180$ .

*Step 8:* Calculate local mean time of event,  $S$ :

$$S = H / 15 + RA / 15 - 0.065710 Y - 6.622$$

Note that  $S$  is expressed in hours and that multiples of 24 should be added or subtracted until  $0 < S < 24$ .

Note that  $S$  is the local time at the point concerned. The corresponding standard time is  $S - B + \beta_m / 15$  h where  $\beta_m$  is the longitude of the standard meridian for the desired time zone (degrees) so that, for example, universal time =  $S - B$ .

## Accuracy of method

The method applies for paths of length 50 to 12 000 km in the LF and MF bands. However, at LF it has only been verified for paths up to 7 500 km.

The special treatment for the southern part of Region 3 might not have been necessary if “corrected geomagnetic latitude” had been used rather than geomagnetic latitude. In any case, the method should be used with caution for geomagnetic latitudes greater than  $60^\circ$ .

Equation (6) which describes how  $G_S$  is modified by the distance  $s_2$  to the next section of land is derived from theory and must therefore be regarded as tentative until measurements are available.

The method predicts the field strength which is likely to be observed if the transmitter and receiver are situated on ground of average conductivity, typically 3 to 10 mS/m. In certain areas (e.g. see Recommendation ITU-R P.832), the effective ground conductivity can be as low as 0.5 mS/m or as high as 40 mS/m. If the ground conductivity at either terminal is an order of magnitude smaller than 10 mS/m, then the field strength may be up to 10 dB smaller. If the ground conductivity at both terminals is an order of magnitude smaller, then the field-strength reduction will be doubled. The amount of attenuation is a function of path length and is greatest for waves approaching grazing incidence. The method may be improved by using a correction for ground conductivity when it differs significantly from that for average ground, for example by using the information contained in the ITU-R Handbook – The ionosphere and its effects on radiowave propagation.

The method assumes that reflection takes place only via the E layer, or that E-layer reflections predominate. However, if  $f > (foE) \sec i$ , where foE is the critical frequency of the E layer and  $i$  is the angle of incidence at the E layer, then the wave will penetrate the E layer and be reflected from the F layer. This is most likely to occur at the highest frequencies in the MF band at ground distances less than 500 km, especially late at night and during the sunspot minimum period. The method can still be used provided  $p$  is calculated for an F-layer reflection height of 220 km and the cymomotive force  $V$  is calculated for the corresponding angle of elevation.

Measurements made in the United States of America suggest that Fig. 3 (hourly loss factor) is likely to be accurate for frequencies near 1 000 kHz in a year of low solar activity. As the frequency deviates in either direction from about 1 000 kHz, particularly during transition hours, appreciable errors may result. These measurements also suggest that the magnitude of the effect of solar activity at two hours after sunset is considerably greater than that at six hours after sunset. Thus, in a year of high solar activity, the difference between field strengths at six hours after sunset and two hours after sunset can be considerably greater than that shown in Fig. 3.

At night, MF sky-waves propagating in temperate latitudes are strongest in spring and autumn and are weakest in summer and winter, the summer minimum being the more pronounced. The overall variation may be as much as 15 dB at the lowest frequencies in the MF band, decreasing to about 3 dB at the upper end of the band. At LF the seasonal variation at night has the opposite trend, with a pronounced summer maximum. The seasonal variation is much smaller in tropical latitudes.

## ANNEX 2

**A discussion on daytime sky-wave propagation****1 LF cases**

Midday field strengths at LF are 7 to 45 dB lower than the values at midnight. The difference is dependent on frequency, distance and season (see also Recommendation ITU-R P.684).

**2 MF cases**

Available data show that midday sky-wave field strengths display a consistent seasonal variation pattern with maximum occurring in winter months. The average winter-month field strength is about 10 dB stronger than the annual median value and the winter-to-summer ratio can exceed 30 dB. The annual median value of midday field strength is about 43 dB lower than its counterpart at six hours after sunset. Field strength exceeded for 10% of the days of the year is about 13 dB stronger than the annual median value. See also the ITU-R Handbook – The ionosphere and its effects on radiowave propagation.

---